Instabilité Hydrodynamique Hydrodynamic Instability

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INTRODUCTION to HYDRODYNAMIC INSTABILITY

- 2019 Stability course 26/9/2019
- Literature
- Some observations of instability
- Equations dimensional analyses & simplications
- Concepts for linear stability analyses
- Various flow Examples of calculations.

Instability course 2019 - Basic approach and equations; normal mode approach. - Shear flow instability in the presence of density differences Kelvin Helmholtz , Hölmböe, Rayleigh-Taylor, Orr Sommerfeld equations Rayleigh and Fjörtöft criterions - Geophysical instabilities: rotational, baroclinic-barotropic, etc, - Centrifugal instabilities: rotational, baroclinic-barotropic, etc, - Convective instabilities Rayleigh Bénard convection (différence de densité) Double diffusion (heat and density diffusion) - Capillary instability (jets, Plateau Rayleigh.) - ... - Interfacial Instabilities (Laurent Davoust) - Magneto HydroDynamic Instability - Nonlinear instability

- *Instabilité Hydrodnamique ; Hydrodynamic instability* Francois Charru, EDP, 2007; En anglais CUP, 2011.

- Introduction to Hydrodynamic Stability Drazin, P.G. & W.H. Reid, Cambridge University Press. (1981) and (2000)

- Hydrodynamic and Hydromagnetic stability Chandrasekhar, S. (1961) Dover

Further advanced reading -Stability and Transition in Shear Flows, P.J. Schmid & D.S. Henningson, Springer, 2001

-Hydrodynamics and Nonlinear instabilities, edited by Godrèche, C. & P. Manneville (1998) Cambridge University Press, Aléa Saclay collection.



Some examples of unstable flows

7



















Instability Mecanism:

Instability: growth of the amplitude of the perturbation of an initially balanced flow

Which balances are there ?

external forces or internal forces

External Forces

- Unstable density distributions (under gravity)
- Centrifugal force
- Coriolis force
- Magneto-Hydro-Dynamic Force
- Surface tension
- ...

Instability Mecanism:

Internal Forces:

- Balance between inertia and pressure force(v=0)
- In shear flows, instability may depend on vorticity dynamics, vortex line stretching and compression.

(Viscous effects often stabilise due to the diffusion of momentum; Definition of Reynolds: $Re=UL/v \approx$ Inertia/viscosity)

The initial state represents a solution of the equations...

EQUATIONS











Simplifications

Boussinesq approximation

$$\frac{\partial \vec{u}}{\partial t} + \vec{u} \cdot \nabla \vec{u} = -\frac{1}{\rho} \nabla p - g \vec{k}$$

For small density variations $\Delta \rho = \rho_2 - \rho_1 \approx 1\%$ we use the Boussinesq approximation, i.e. only density variations in z are considered

The inertia effect on density, i.e. $\Delta \rho \partial u / \partial t$ is neglected and only the effect of the gravitational acceleration

$$g \Delta \rho$$
 (or often $g'=g \Delta \rho/\rho_{mean}$)
on $\Delta \rho$ taken into account.

Hydrostatic balance

If the aspect ratio H/L is small, than we have for horizontal density perturbations, to leading order $\frac{\partial p}{\partial z} = -\rho g$

25

Bernoulli equation Suppose a homogeneous, barotropic flow, no density effects, and neglect viscous effect (v=0) so that we have the Euler equations: $\frac{\partial \vec{u}}{\partial t} + \vec{u} \cdot \nabla \vec{u} = -\frac{1}{o} \nabla p - g \vec{k}$ $\nabla \vec{u} = 0$ Introduce a gravitational potential $\Phi_{\rm gr}$ with $gk = grad \Phi_{\rm gr}$ so that $\Phi_{\rm gr} = -gz$ For the nonlinear term we use the vector identity $(\vec{u}.\nabla)\vec{u} = \frac{1}{2}\nabla(\vec{u}.\vec{u}) + (\nabla \times \vec{u}) \times \vec{u}$ and obtain for the Euler equation $\frac{\partial \vec{u}}{\partial t} + \vec{\omega} \times \vec{u} + \nabla (\frac{1}{2}U^2) - \nabla \Phi_{gr} + \frac{1}{\rho} \nabla p = 0$ vorticity = $\vec{\omega} \equiv \nabla \times \vec{u}$ $U^2 = |\vec{u}.\vec{u}|$ With p=p(ρ) we may write: $\frac{\partial \vec{u}}{\partial t} + \vec{\omega} \times \vec{u} = -\nabla \left(\frac{1}{2}U^2 - \Phi_{gr} + \int \frac{\nabla p}{\rho}\right) = \nabla H$ (A)

H is a scalar potential function. We consider a few cases:

1) Steady flow: $\vec{\omega} \times \vec{u} = -\nabla H$ Since $\vec{u} \times (\nabla \times \vec{u}) = \vec{u} \times \vec{\omega} \equiv 0$ along a streamline, we obtain: $\frac{1}{2}U^2 + gz + \int \frac{\nabla p}{2} = H = \text{constant along streamlines}$ 2) irrotational : $\omega = 0$ we can introduce the velocity potential $\vec{u} = \nabla \phi$ $\frac{\partial \phi}{\partial t} + \frac{1}{2}U^2 + gz + \int \frac{\nabla p}{\rho} = f(t)$ in the entire flow field *f(t)* is a function of time, and U^2 can be written as $\nabla \phi \cdot \nabla \phi$ 3) Steady & irrotational flow: $\frac{1}{2}U^2 + gz + \int \frac{\nabla p}{2} = H = \text{constant in the entire flow field}$ equations 3) and (A) above are known as the Bernoulli equation !

4) Steady flow with H= constant: The Euler equation becomes: $\omega \times \vec{u} = 0$ In 2D flows this implies: $\omega = 0$ In 3D flows ω is parallel to \vec{u} These are known as Beltrami flows.

PERTURBATION OF EQUATIONS AND LINEARIZATION

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In linear stability analyses, one supposes a steady basic state U_0 that is perturbed with a perturbation v', which is restricted to be infinitesimal. The precise meaning of 'infinitesimal' depends on the physical context and the particular experiment.

For instance, one may expand its amplitude A in a Taylor series :

$$A = A_0 + \epsilon A_1 + \epsilon^2 A_2 + \dots$$

where A_0 is the amplitude of the basic flow, and the small parameter ϵ is a small number that is characteristic for the system under consideration. For instance, when the Reynolds number Re >> 1 characterizes the flow, ϵ can be chosen as $\epsilon = 1/Re$.

To know which numbers do characterize the flow we may use, e.g. physical arguments or dimensional analyses.

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The perturbation equations are obtained after inserting the time-dependent perturbation. With v of order ϵ we consider leading and first order, i.e. the linear approximation :

$$\mathbf{u}(\mathbf{x},t) = \mathbf{U}_0(\mathbf{x}) + \mathbf{v}(\mathbf{x},t),$$

$$P(\mathbf{x},t) = P_0(\mathbf{x}) + p(\mathbf{x},t)$$

in the equations of motion.

For the Euler equations (viscosity $\nu = 0$), the steady fields must satisfy

 $(\mathbf{U}_0 \cdot \nabla)\mathbf{U}_0 = -\nabla P_0$ $\mathbf{?}$ $\nabla \cdot \mathbf{U}_0 = 0.$

and continuity

Linearization and making use of the basic state U_0 gives

$$\frac{\partial \mathbf{v}}{\partial t} = -(\mathbf{U}_0 \cdot \nabla)\mathbf{v} - (\mathbf{v} \cdot \nabla) \cdot \mathbf{U}_0 - \nabla p \tag{1}$$

and

$$abla \cdot \mathbf{v} = \mathbf{0}$$

with initial conditions $\mathbf{v}(\mathbf{x}, 0) = \mathbf{v}_0(\mathbf{x})$ and boundary conditions. Subsequently we choose a perturbation amplitude in the form of periodic waves.

(Note that the basic state is subtracted from the equations to obtain the perturbation equations)

Time is eliminated by the Laplace transform of the system with respect to t, seeking solutions of the form

$$\hat{v}_k(r,t) = \hat{v}_k(r)e^{s_kt}$$

where $s_k = s(k) = \omega_R(k) + i\omega_i(k)$ is a complex constant to be determined with the stability analyses; its value may be different for each different k.

The velocity \hat{v} is to be found from the initial basic velocity field, and the transformed system of ordinary differential equations in r, and the boundary conditions in r.

The perturbation function

The choice of the perturbation function depends on the flow geometry and initial conditions. For a system limited in z and open in x-direction, the perturbation is

 $\sim \hat{v}(z)e^{i(kx+s_kt)}$

(e.g. Kelvin-Helmholtz $\hat{v}(z)$ is determined with $\nabla^2 v = 0$); For a Poiseuille flow in r-direction and open in x we must analyse perturbations of the form

$$\sim \hat{v}(r)e^{i(s_kt+kx+n\theta)}$$
 and $s_k = \omega_R + i\omega_i$

Derivatives in x and θ in the equations of motion are transformed into *ik* and *in*, repectively, whereas differentiation in the *r*-direction, ∂_r , leads to an ordinary differential equation in *r* that needs to be solved with the boundary conditions.

Other boundary geometries

For problems with a spherical geometry one would take

$$\hat{v}(x,r,\theta,t) = \sum_{l=0}^{\infty} \sum_{m=-l}^{m=+l} \hat{v}_l^m(r,t) Y_l^m(\theta,\phi)$$

where $Y_I^m(\theta, \phi)$ represent spherical harmonics. Now, the behaviour of the system with respect to modes *I* and *m* has to be investigated. Thus, in all cases the disturbance is expanded in a suitable set of normal modes in accordance with flow geometry.

Legendre functions

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Other boundary geometries

The choice of the symmetry of the disturbance depends on the geometry of the system. For flows with a symmetry axis, for example in the case of a Poiseuille flow, one would take

$$\hat{v}(x, r, \theta, t) = \sum_{n=-\infty}^{\infty} \int_{-\infty}^{\infty} \hat{v}_{k,n}(r, t) e^{ikx+in\theta} dk$$

with p, v etc. functions of r. We analyse an arbitrary function in terms of two-dimensional periodic waves with amplitude $\hat{v}(x, r, \theta)$ where $k = \sqrt{k^2 + n^2}$ is the wave number associated with the disturbance $\hat{v}_{k,n}$.

Some concepts

The system of the Hagen-Poiseuille flow $U = V(1 - z^2/d^2)\bar{e}_x$ takes the form of an eigenvalue relation

$$F(s,k,n,V,d,\nu)=0$$

and eigenfunctions \hat{v} , p. (ν is the kinematic viscosity $\nu = \mu/\rho$). This so-called *method of normal modes* makes use of small disturbances that are resolved into modes which satisfy the linear system and therefore may be treated separately. The use of the Laplace Fourrier transform, thus reduces the equations of motion to an ordinary equation or even an algebraic equation in the parameters of F.

The dispersion relation

Some concepts

The solution of the ODE (or PDE) + boundary conditions provides the *dispersion relation* for *s*

 $s = s_n(R, k)$

where k is the wave number and R the set of control parameters. such as for instance the Reynolds number in the case of the Poiseuille flow.

The fastest growing mode k_c appears the first, and the critical value above which instability occurs.

Because the system is linear, the real and imaginary parts are separate solutions. For stability analyses we are generally interested in the real part of the solutions i.e.

$$\hat{v}_k(r,t) = Re\{\hat{v}_k(r)e^{s_kt}\}$$

without explicitly mentioning it.

The dispersion relation, and stability interpretation Some concepts

The growth rate :

(suppose $s = \omega_R + i\omega_i$ and perturbations $\sim e^{st}$ the real part ω_R the exponential growth and the imaginary part ω_i , the sinusoidal part).

for $\omega_R < 0$ the flow is stable for $\omega_R = 0$ the flow is neutrally stable . for $\omega_R > 0$ there is exponential growth.

A flow is marginally stable when $\omega_R = 0$ for critical values on which the eigenvalue ω_R depends, but $\omega_R > 0$ for some neighbouring values of the parameters. On a neutral curve $\omega_R = 0$, but ω_R is not positive for any of the neighbouring parameters.

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EXAMPLE 1

class exercise : surface waves perturbation and linearisation:

$$\frac{\rho_1 << \rho_0 \rightarrow \rho_1 \approx 0}{\rho_0}$$

hydrostatic balance $\frac{dp_0}{dz}$

$$\frac{d}{dt} = -\rho g$$

- Consider a basin at rest in hydrostatic balance $\boldsymbol{u}'=(\boldsymbol{u}', \boldsymbol{w}')$ whereas $p=p_0(\boldsymbol{x},\boldsymbol{z})+p'$ and $\rho=\rho_0$.
- Use Bernouilli Give the expressions for the leading order O(1) balance and second order $O(\varepsilon)$ balance.
- Derive the dispersion relation

TWO EXAMPLES

Surface waves

Kelvin Helmholtz





Velocity potential in each layer $u = \nabla \phi$ $u = \frac{\partial \phi}{\partial x}$ $w = \frac{\partial \phi}{\partial z}$ Continuity $\nabla . u = 0$ $\nabla^2 \phi_i = 0$ Laplace equation Note below: upper layer has index *i*=1, lower layer has index *i*=2

Potential flow above and below the interface, we may use *Bernoulli*, $\omega = 0$

$$\frac{\partial \phi_i}{\partial t} + \frac{1}{2} \left(\nabla \phi_i \right)^2 + gz + \frac{P_i}{\rho} = 0$$





The perturbations	Basic flow + perturbation						
above the interface $\phi_1 = \phi_{1b} + \epsilon \phi'_1 + \dots$ with $\phi_{1b} = -\frac{1}{2}Ux$							
below the interface $\phi_2 = \phi_{2b} + \epsilon \phi'_2 + \dots$ with $\phi_{1b} = -\frac{1}{2}Ux$							
Substitute in the kinemati	c interface condition	ı					
$w_i = \frac{D\zeta}{Dt} = \frac{\partial\zeta}{\partial t}$	$+ \frac{\partial \phi_i}{\partial x} \frac{\partial \zeta}{\partial x} + \frac{\partial \phi_i}{\partial z} + \frac{\partial \phi_i}{\partial z} \frac{\partial \phi_i}{\partial z} \frac{\partial \phi_i}{\partial z} \frac{\partial \phi_i}{\partial z} + \frac{\partial \phi_i}{\partial z} \frac{\partial \phi_i}{\partial z} \frac{\partial \phi_i}{\partial z} + \frac{\partial \phi_i}{\partial z} \frac{\partial \phi_i}{\partial z} \frac{\partial \phi_i}{\partial z} + \frac{\partial \phi_i}{\partial z} \frac{\partial \phi_i}{\partial z} \frac{\partial \phi_i}{\partial z} + \frac{\partial \phi_i}{\partial z} \frac{\partial \phi_i}{\partial z} \frac{\partial \phi_i}{\partial z} + \frac{\partial \phi_i}{\partial z} \frac{\partial \phi_i}{\partial z} \frac{\partial \phi_i}{\partial z} + \frac{\partial \phi_i}{\partial z} \frac{\partial \phi_i}{\partial z} \frac{\partial \phi_i}{\partial z} + \frac{\partial \phi_i}{\partial z} \frac{\partial \phi_i}{\partial z} \frac{\partial \phi_i}{\partial z} + \partial $	$\frac{\partial \zeta}{\partial z}$ $\zeta = O(\epsilon)$					
$w_1 = \frac{\partial \phi_1'}{\partial z} = \frac{D\zeta}{Dt} =$	$= \frac{\partial \zeta}{\partial t} + \left(-\frac{1}{2}U + \frac{\partial \phi_1'}{\partial x}\right)$	$\bigg)_{z=\zeta} \frac{\partial \zeta}{\partial x} + \frac{\partial \phi_1'}{\partial z} \frac{\partial \zeta}{\partial z}$					
$w_2 = \frac{\partial \phi_2'}{\partial z} = \frac{D\zeta}{Dt} =$	$=\frac{\partial\zeta}{\partial t}+\left(\frac{1}{2}U+\frac{\partial\phi_2'}{\partial x}\right)$	$\int_{z=\zeta} \frac{\partial \zeta}{\partial x} + \frac{\partial \phi_2'}{\partial z} \frac{\partial \zeta}{\partial z}$					
Linear approximation : ke	eep O(ε) terms,	$w_1 = \frac{\partial \phi_1'}{\partial z} = \frac{\partial \zeta}{\partial t} - \frac{1}{2}U\frac{\partial \zeta}{\partial x}$					
ne	eglect O(ε²) terms:	$w_2 = \frac{\partial \phi_2'}{\partial z} = \frac{\partial \zeta}{\partial t} + \frac{1}{2}U\frac{\partial \zeta}{\partial x}$					

In dynamic boundary condition

$$z = \zeta \qquad \frac{\partial \phi_1}{\partial t} + \frac{1}{2} \left(\nabla \phi_1 \right)^2 = \frac{\partial \phi_2}{\partial t} + \frac{1}{2} \left(\nabla \phi_2 \right)^2$$
with $\nabla \phi_1 = -\frac{1}{2}U + \frac{\partial \phi'_1}{\partial x}$ and $\nabla \phi_2 = \frac{1}{2}U + \frac{\partial \phi'_2}{\partial x}$
 $\left(\frac{\partial \phi_2}{\partial t} - \frac{\partial \phi_1}{\partial t} \right)_{z=\zeta} = \frac{U}{2} \left(\frac{\partial \phi'_2}{\partial t} + \frac{\partial \phi'_1}{\partial t} \right)_{z=\zeta}$

From the three	conditions we have:
Laplace equatio	$\mathbf{n} \nabla^2 \phi_i = 0$
Kinematic BC	$w_1 = \frac{\partial \phi'_1}{\partial z} = \frac{\partial \zeta}{\partial t} - \frac{1}{2}U\frac{\partial \zeta}{\partial x}$ $w_2 = \frac{\partial \phi'_2}{\partial z} = \frac{\partial \zeta}{\partial t} + \frac{1}{2}U\frac{\partial \zeta}{\partial x}$
Dynamic BC	$\left(\frac{\partial\phi_2}{\partial t} - \frac{\partial\phi_1}{\partial t}\right)_{z=\zeta} = \frac{U}{2} \left(\frac{\partial\phi_2'}{\partial t} + \frac{\partial\phi_1'}{\partial t}\right)_{z=\zeta}$

perturbations:

$$\phi_i = F(z) \exp(ikx + \sigma t)$$
 and $\zeta = A \exp(ikx + \sigma t)$

With F(z) the vertical dependence to determine, and A the amplitude

The wave form is sinusoidal, with spacing $\lambda = 2\pi/k$

With the Laplace transform, $e^{\sigma t}$ exponential decay or growth is supposed

$$\nabla^2 \phi_i = 0 \quad \frac{d^2 \phi_i}{dx^2} + \frac{d^2 \phi_i}{dz^2} = 0$$
$$\frac{d^2 F}{dz^2} - k^2 F = 0$$

Condition at infinity: the amplitude of the perturbations goes to zero

$$\begin{split} \phi_i &= B_1 e^{-kz} + B_2 e^{kz} \\ \phi_i &\to 0 \qquad z \to +\infty \text{ thus for } z > 0 \ B_2 = 0 \\ \phi_i &\to 0 \qquad z \to -\infty \text{ thus for } z < 0 \ B_1 = 0 \end{split}$$

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We can now solve the form of \zeta^*, \varphi^{*_1}, \varphi^{*_2} with amplitudes A, B<sub>1</sub>, and B<sub>2</sub>

\zeta = A e^{ikx + \sigma t}.

\varphi'_1 = B_1 e^{-kz} e^{ikx + \sigma t}. \varphi'_2 = B_2 e^{kz} e^{ikx + \sigma t}

Substitution in conditions 1 and II:

-kB_1 = (\sigma - \frac{1}{2}i \ k \ U) A

-kB_2 = (\sigma + \frac{1}{2}i \ k \ U) A

and condition III: i [\sigma(B_2 - B_1)_{z=0} + \frac{1}{2} U (B_2 k + B_1 k)_{z=0}] e^{i(kx)} = 0

\sigma = \frac{1}{2} ik(U_1 + U_2) \pm \frac{1}{2} k(U_1 - U_2)

for U_1 = -U_2 this reduces to

\sigma = \pm kU

\sigma(k) is the dispersion relation showing the variation of growth rate with k.

For \sigma > 0, k \neq 0 the sheet is unstable. Small wavelengths grow faster than short ones.
```







Scaling

For this Poisseuille flow the parameter *a* to determine is the pressure gradient

$$\frac{dp}{dx} = f(U, d, \rho, \mu)$$

we have the dimensions :

$$\begin{bmatrix} \frac{dp}{dx} \end{bmatrix} = ML^{-2}T^{-2}$$
$$\begin{bmatrix} U \end{bmatrix} = LT^{-1}$$
$$\begin{bmatrix} D \end{bmatrix} = L$$
$$\begin{bmatrix} \rho \end{bmatrix} = ML^{-3}$$
$$\begin{bmatrix} \mu \end{bmatrix} = ML^{-1}T^{-1}$$

In this case U, D, and ρ are independent $([\mu] = [\rho][u][D])$, so that k = 3, m = 1 and n = 4 and so $\Pi = \Phi(\Pi_1)$

Scaling

Buckingham theorem: find dimensionless numbers

Consider the non-dimensional parameter a to determine, that depends on n governing parameters

 $a = f(a_1, a_2, ..., a_k, b_1, ..., b_m)$

with k independent parameters a_i , and m dependent parameters b_i , m + k = n and f is a function. Since b_i are dependent parameters we can write them as a function of the independent parameters a_i :

 $b_1 = [a_1]^{p_1} \dots [a_k]^{r_1}$

$$b_m = [a_1]^{p_m} ... [a_k]'$$

Scaling

n=4 governing parameters with V, D, ρ independent and one dependent parameters μ . So

$$\Pi_1 = \frac{\mu}{UD\rho}$$

(this is 1/Re) and Π the dimensionless pressure gradient

$$\exists = \frac{1}{(U^2 D^{-1} \rho)} \frac{d\rho}{dx}$$

We thus obtain :

$$\Pi = (U^2 D^{-1} \rho)^{-1} \frac{dp}{dx} = \Phi(\Pi_1) = \Phi(1/Re)$$

and Φ the function to determine. Whatever the individual values of d, U, μ or ρ are, this function is universal. Here Φ is determined experimentally, see picture.

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The characteristic times of diffusion are:

Intro

 $\tau_v = h^2/v \dots$ for vorticity diffusion $\tau_{\theta} = h^2/\kappa \dots$ for heat diffusion

There is a competition between the acceleration (τ_A) opposed by diffusion effects. The ratio in time scales determines stability:

Ra= $\tau_v \tau_{\theta} / \tau_A^2 = \frac{\alpha g \Delta T h^3}{\kappa v}$ (Rayleigh number) Ra>1 convection Ra<1 stable... in reality Ra>673 or higher (but the number is right)







Home exercise: A NORMAL MODE TOY PROBLEM Suppose there is a balance represented by the 1D equation:

 $\frac{\partial f(y)}{\partial t} = f(y) - f(y)^2 + \frac{1}{\lambda} \frac{\partial^2 f(y)}{\partial y^2}$

with basic state f=0, and boundary conditions f(0)=f(1)=0.

Is the basic state a solution of the equation ?
 Perturb the basic flow by adding a perturbation *f*'

$f(x,t) = \bar{f} + \epsilon f'(x,t) + \epsilon^2 \dots$

3. Substitute, separate O(0), O(ϵ), ...and consider order ϵ only 4. What happens at O(0) ? Consider the equation for O(ϵ).

5. Use the Laplace transformation $f'(t) = F(y) e^{\sigma(k)t}$ and solve F(y) with the boundary conditions.

6. What is the dispersion relation and what does this relation show ?